# Appendix D

# **Probability Distributions**

This appendix archives a number of useful results from texts by Papoulis [44], Lee [33] and Cover [12]. Table 16.1 in Cover (page 486) gives entropies of many distributions not listed here.

## **D.1** Transforming PDFs

Because probabilities are defined as areas under PDFs when we transform a variable

$$y = f(x) \tag{D.1}$$

we transform the PDF by preserving the areas

$$p(y)|dy| = p(x)|dx| \tag{D.2}$$

where the absolute value is taken because the changes in x or y (dx and dy) may be negative and areas must be positive. Hence

$$p(y) = \frac{p(x)}{\left|\frac{dy}{dx}\right|} \tag{D.3}$$

where the derivative is evaluated at  $x = f^{-1}(y)$ . This means that the function f(x) must be one-to-one and invertible.

If the function is many-to-one then it's inverse will have multiple solutions  $x_1, x_2, ..., x_n$ and the PDF is transformed at each of these points (Papoulis' Fundamental Theorem [44], page 93)

$$p(y) = \frac{p(x_1)}{\left|\frac{dy}{dx_1}\right|} + \frac{p(x_2)}{\left|\frac{dy}{dx_2}\right|} + \dots + \frac{p(x_n)}{\left|\frac{dy}{dx_n}\right|}$$
(D.4)

#### D.1.1 Mean and Variance

For more on the mean and variance of functions of random variables see Weisberg [64].

Expectation is a *linear operator*. That is

$$E[(a_1x + a_2x)] = a_1E[x] + a_2E[x]$$
(D.5)

Therefore, given the function

$$y = ax \tag{D.6}$$

we can calculate the mean and variance of y as functions of the mean and variance of x.

$$E[y] = aE[x]$$

$$Var(y) = a^{2}Var(x)$$
(D.7)

If y is a function of many *uncorrelated* variables

$$y = \sum_{i} a_i x_i \tag{D.8}$$

we can use the results

$$E[y] = \sum_{i} a_i E[x_i] \tag{D.9}$$

$$Var[y] = \sum_{i} a_i^2 Var[x_i]$$
 (D.10)

But if the variables are correlated then

$$Var[y] = \sum_{i} a_{i}^{2} Var[x_{i}] + 2 \sum_{i} \sum_{j} a_{i}a_{j} Var(x_{i}, x_{j})$$
(D.11)

where  $Var(x_i, x_j)$  denotes the covariance of the random variables  $x_i$  and  $x_j$ .

#### **Standard Error**

As an example, the mean

$$m = \frac{1}{N} \sum_{i} x_i \tag{D.12}$$

of uncorrelated variables  $x_i$  has a variance

$$\sigma_m^2 \equiv Var(m) = \sum_i \frac{1}{N} Var(x_i)$$

$$= \frac{\sigma_x^2}{N}$$
(D.13)

where we have used the substitution  $a_i = 1/N$  in equation D.10. Hence

$$\sigma_m = \frac{\sigma_x}{\sqrt{N}} \tag{D.14}$$

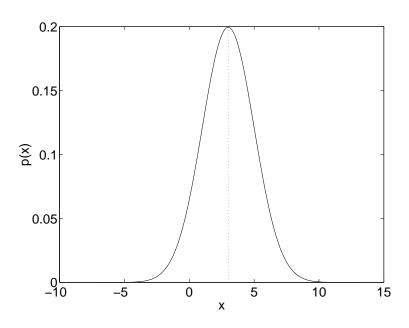


Figure D.1: The Gaussian Probability Density Function with  $\mu = 3$  and  $\sigma = 2$ .

# D.2 Uniform Distribution

The uniform PDF is given by

$$U(x;a,b) = \frac{1}{b-a} \tag{D.15}$$

for  $a \le x \le b$  and zero otherwise. The mean is 0.5(a+b) and variance is  $(b-a)^2/12$ . The entropy of a uniform distribution is

$$H(x) = \log(b - a) \tag{D.16}$$

# D.3 Gaussian Distribution

The Normal or Gaussian probability density function, for the case of a single variable, is

$$N(x;\mu,\sigma^2) = \frac{1}{(2\pi\sigma^2)^{1/2}} \exp\left(-\frac{(x-\mu)^2}{2\sigma^2}\right)$$
(D.17)

where  $\mu$  and  $\sigma^2$  are the mean and variance.

### D.3.1 Entropy

The entropy of a Gaussian variable is

$$H(x) = \frac{1}{2}\log\sigma^2 + \frac{1}{2}\log 2\pi + \frac{1}{2}$$
(D.18)

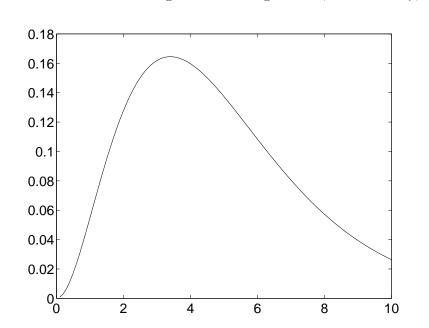


Figure D.2: The Gamma Density for b = 1.6 and c = 3.125.

For a given variance, the Gaussian distribution has the highest entropy. For a proof of this see Bishop ([3], page 240).

### D.3.2 Relative Entropy

For Normal densities  $q(x) = N(x; \mu_q, \sigma_q^2)$  and  $p(x) = N(x; \mu_p, \sigma_p^2)$  the KL-divergence is

$$D[q||p] = \frac{1}{2}\log\frac{\sigma_p^2}{\sigma_q^2} + \frac{\mu_q^2 + \mu_p^2 + \sigma_q^2 - 2\mu_q\mu_p}{2\sigma_p^2} - \frac{1}{2}$$
(D.19)

# D.4 The Gamma distribution

The Gamma density is defined as

$$\Gamma(x;b,c) = \frac{1}{\Gamma(c)} \frac{x^{c-1}}{b^c} \exp\left(\frac{-x}{b}\right)$$
(D.20)

where  $\Gamma()$  is the gamma function [49]. The mean of a Gamma density is given by bcand the variance by  $b^2c$ . Logs of gamma densities can be written as

$$\log \Gamma(x; b, c) = \frac{-x}{b} + (c - 1) \log x + K$$
 (D.21)

where K is a quantity which does not depend on x; the log of a gamma density comprises a term in x and a term in  $\log x$ . The Gamma distribution is only defined for positive variables.

#### D.4.1 Entropy

Using the result for Gamma densities

$$\int p(x) \log x = \Psi(c) + \log b \tag{D.22}$$

where  $\Psi()$  is the digamma function [49] the entropy can be derived as

$$H(x) = \log\Gamma(c) + c\log b - (c-1)(\Psi(c) + \log b) + c$$
 (D.23)

#### D.4.2 Relative Entropy

For Gamma densities  $q(x) = \Gamma(\boldsymbol{x}; b_q, c_q)$  and  $p(x) = \Gamma(\boldsymbol{x}; b_p, c_p)$  the KL-divergence is

$$D[q||p] = (c_q - 1)\Psi(c_q) - \log b_q - c_q - \log \Gamma(c_q)$$

$$+ \log \Gamma(c_p) + c_p \log b_p - (c_p - 1)(\Psi(c_q) + \log b_q) + \frac{b_q c_q}{b_p}$$
(D.24)

# **D.5** The $\chi^2$ -distribution

If  $z_1, z_2, ..., z_N$  are independent normally distributed random variables with zero-mean and unit variance then

$$x = \sum_{i=1}^{N} z_i^2$$
 (D.25)

has a  $\chi^2$ -distribution with N degrees of freedom ([33], page 276). This distribution is a special case of the Gamma distribution with b = 2 and c = N/2. This gives

$$\chi^{2}(x;N) = \frac{1}{\Gamma(N/2)} \frac{x^{N/2-1}}{2^{N/2}} \exp\left(\frac{-x}{2}\right)$$
(D.26)

The mean and variance are N and 2N. The entropy and relative entropy can be found by substituting the the values b = 2 and c = N/2 into equations D.23 and D.24. The  $\chi^2$  distribution is only defined for positive variables.

If x is a  $\chi^2$  variable with N degrees of freedom and

$$y = \sqrt{x} \tag{D.27}$$

then y has a  $\chi$ -density with N degrees of freedom. For N = 3 we have a Maxwell density and for N = 2 a Rayleigh density ([44], page 96).

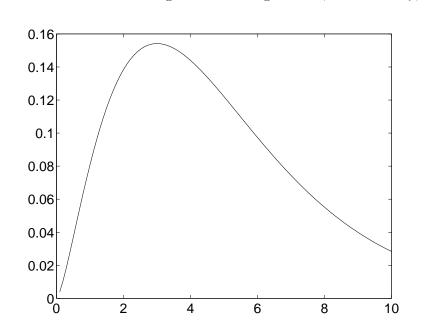


Figure D.3: The  $\chi^2$  Density for N = 5 degrees of freedom.

# D.6 The t-distribution

If  $z_1, z_2, ..., z_N$  are independent Normally distributed random variables with mean  $\mu$ and variance  $\sigma^2$  and m is the sample mean and s is the sample standard deviation then

$$x = \frac{m - \mu}{s / \sqrt{N}} \tag{D.28}$$

has a t-distribution with N-1 degrees of freedom. It is written

$$t(x;D) = \frac{1}{B(D/2,1/2)} \left(1 + \frac{x^2}{D}\right)^{-(D+1)/2}$$
(D.29)

where D is the number of 'degrees of freedom' and

$$B(a,b) = \frac{\Gamma(a)\Gamma(b)}{\Gamma(a+b)}$$
(D.30)

is the *beta function*. For D = 1 the t-distribution reduces to the standard Cauchy distribution ([33], page 281).

# D.7 Generalised Exponential Densities

The 'exponential power' or 'generalised exponential' probability density is defined as

$$p(a) = G(a; R, \beta) = \frac{R\beta^{1/R}}{2\Gamma(1/R)} \exp(-\beta |a|^R)$$
(D.31)

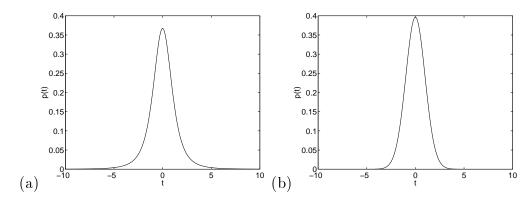


Figure D.4: The t-distribution with (a) N = 3 and (b) N = 49 degrees of freedom.

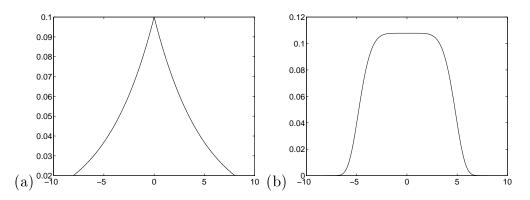


Figure D.5: The generalised exponential distribution with (a) R = 1, w = 5 and (b) R = 6, w = 5. The parameter R fixes the weight of the tails and w fixes the width of the distribution. For (a) we have a Laplacian which has positive kurtosis (k = 3); heavy tails. For (b) we have a light-tailed distribution with negative kurtosis (k = -1).

where  $\Gamma()$  is the gamma function [49], the mean of the distribution is zero<sup>1</sup>, the width of the distribution is determined by  $1/\beta$  and the weight of its tails is set by R. This gives rise to a Gaussian distribution for R = 2, a Laplacian for R = 1 and a uniform distribution in the limit  $R \to \infty$ . The density is equivalently parameterised by a variable w, which defines the width of the distribution, where  $w = \beta^{-1/R}$  giving

$$p(a) = \frac{R}{2w\Gamma(1/R)} \exp(-|a/w|^R)$$
 (D.32)

The variance is

$$V = w^2 \frac{\Gamma(3/R)}{\Gamma(1/R)} \tag{D.33}$$

which for R = 2 gives  $V = 0.5w^2$ . The kurtosis is given by [7]

$$K = \frac{\Gamma(5/R)\Gamma(1/R)}{\Gamma(3/R)^2} - 3$$
(D.34)

where we have subtracted 3 so that a Gaussian has zero kurtosis. Samples may be generated from the density using a rejection method [59].

<sup>&</sup>lt;sup>1</sup>For non zero mean we simply replace a with  $a - \mu$  where  $\mu$  is the mean.

# D.8 PDFs for Time Series

Given a signal a = f(t) which is sampled uniformly over a time period T, its PDF, p(a) can be calculated as follows. Because the signal is uniformly sampled we have p(t) = 1/T. The function f(t) acts to transform this density from one over t to to one over a. Hence, using the method for transforming PDFs, we get

$$p(a) = \frac{p(t)}{\left|\frac{da}{dt}\right|} \tag{D.35}$$

where || denotes the absolute value and the derivative is evaluated at  $t = f^{-1}(x)$ .

#### D.8.1 Sampling

When we convert an analogue signal into a digital one the sampling process can have a crucial effect on the resulting density. If, for example, we attempt to sample uniformly but the sampling frequency is a multiple of the signal frequency we are, in effect, sampling non-uniformly. For true uniform sampling it is necessary that the ratio of the sampling and signal frequencies be irrational.

#### D.8.2 Sine Wave

For a sine wave,  $a = \sin(t)$ , we get

$$p(a) = \frac{1}{|cos(t)|} \tag{D.36}$$

where cos(t) is evaluated at  $t = \sin^{-1}(a)$ . The inverse sine is only defined for  $-\pi/2 \le t \le \pi/2$  and p(t) is uniform within this. Hence,  $p(t) = 1/\pi$ . Therefore

$$p(a) = \frac{1}{\pi\sqrt{1-a^2}}$$
(D.37)

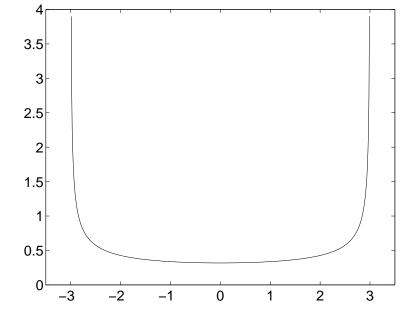
This density is *multimodal*, having peaks at +1 and -1. For a more general sine wave

$$a = R\sin(wt) \tag{D.38}$$

we get  $p(t) = w/\pi$ 

$$p(a) = \frac{1}{\pi\sqrt{1 - (a/R)^2}}$$
(D.39)

which has peaks at  $\pm R$ .



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Figure D.6: The PDF of  $a = R \sin(wt)$  for R = 3.

# Appendix E

# Multivariate Probability Distributions

## E.1 Transforming PDFs

Just as univariate Probability Density Functions (PDFs) are transformed so as to preserve area so multivariate probability distributions are transformed so as to preserve volume. If

$$\boldsymbol{y} = f(\boldsymbol{x}) \tag{E.1}$$

then this can be achieved from

$$p(\boldsymbol{y}) = \frac{p(\boldsymbol{x})}{abs(|\boldsymbol{J}|)}$$
(E.2)

where abs() denotes the absolute value and || the determinant and

$$\boldsymbol{J} = \begin{bmatrix} \frac{\partial y_1}{\partial x_1} & \frac{\partial y_1}{\partial x_2} & \cdots & \frac{\partial y_1}{\partial x_d} \\ \frac{\partial y_2}{\partial x_1} & \frac{\partial y_2}{\partial x_2} & \cdots & \frac{\partial y_2}{\partial x_d} \\ \cdots & \cdots & \cdots & \cdots \\ \frac{\partial y_d}{\partial x_1} & \frac{\partial y_d}{\partial x_2} & \cdots & \frac{\partial y_d}{\partial x_d} \end{bmatrix}$$
(E.3)

is the Jacobian matrix for d-dimensional vectors  $\boldsymbol{x}$  and  $\boldsymbol{y}$ . The partial derivatives are evaluated at  $\boldsymbol{x} = f^{-1}(\boldsymbol{y})$ . As the determinant of  $\boldsymbol{J}$  measures the volume of the transformation, using it as a normalising term therefore preserves the volume under the PDF as desired. See Papoulis [44] for more details.

#### E.1.1 Mean and Covariance

For a vector of random variables (Gaussian or otherwise),  $\boldsymbol{x}$ , with mean  $\boldsymbol{\mu}_x$  and covariance  $\boldsymbol{\Sigma}_x$  a linear transformation

$$\boldsymbol{y} = \boldsymbol{F}\boldsymbol{x} + \boldsymbol{C} \tag{E.4}$$

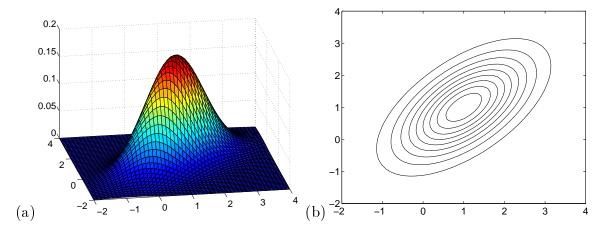


Figure E.1: (a) 3D-plot and (b) contour plot of Multivariate Gaussian PDF with  $\boldsymbol{\mu} = [1, 1]^T$  and  $\boldsymbol{\Sigma}_{11} = \boldsymbol{\Sigma}_{22} = 1$  and  $\boldsymbol{\Sigma}_{12} = \boldsymbol{\Sigma}_{21} = 0.6$  ie. a positive correlation of r = 0.6.

gives rise to a random vector  $\boldsymbol{y}$  with mean

$$\boldsymbol{\mu}_y = \boldsymbol{F} \boldsymbol{\mu}_x + \boldsymbol{C} \tag{E.5}$$

and covariance

$$\boldsymbol{\Sigma}_y = \boldsymbol{F} \boldsymbol{\Sigma}_x \boldsymbol{F}^T \tag{E.6}$$

If we generate another random vector, this time from a  $different\, {\rm linear}\, {\rm transformation}$  of  ${\pmb x}$ 

$$\boldsymbol{z} = \boldsymbol{G}\boldsymbol{x} + \boldsymbol{D} \tag{E.7}$$

then the covariance *between* the random vectors  $\boldsymbol{y}$  and  $\boldsymbol{z}$  is given by

$$\boldsymbol{\Sigma}_{y,z} = \boldsymbol{F} \boldsymbol{\Sigma}_x \boldsymbol{G}^T \tag{E.8}$$

The i,jth entry in this matrix is the covariance between  $y_i$  and  $z_j$ .

## E.2 The Multivariate Gaussian

The multivariate normal PDF for d variables is

$$N(\boldsymbol{x};\boldsymbol{\mu},\boldsymbol{\Sigma}) = \frac{1}{(2\pi)^{d/2} |\boldsymbol{\Sigma}|^{1/2}} \exp\left(-\frac{1}{2}(\boldsymbol{x}-\boldsymbol{\mu})^T \boldsymbol{\Sigma}^{-1}(\boldsymbol{x}-\boldsymbol{\mu})\right)$$
(E.9)

where the mean  $\boldsymbol{\mu}$  is a d-dimensional vector,  $\boldsymbol{\Sigma}$  is a  $d \times d$  covariance matrix, and  $|\boldsymbol{\Sigma}|$  denotes the determinant of  $\boldsymbol{\Sigma}$ .

#### E.2.1 Entropy

The entropy is

$$H(\boldsymbol{x}) = \frac{1}{2}\log|\boldsymbol{\Sigma}| + \frac{d}{2}\log 2\pi + \frac{d}{2}$$
(E.10)

#### E.2.2 Relative Entropy

For Normal densities  $q(\boldsymbol{x}) = N(\boldsymbol{x}; \boldsymbol{\mu}_q, \boldsymbol{\Sigma}_q)$  and  $p(\boldsymbol{x}) = N(\boldsymbol{x}; \boldsymbol{\mu}_p, \boldsymbol{\Sigma}_p)$  the KL-divergence is

$$D[q||p] = 0.5 \log \frac{|\boldsymbol{\Sigma}_p|}{|\boldsymbol{\Sigma}_q|} + 0.5 Trace(\boldsymbol{\Sigma}_p^{-1}\boldsymbol{\Sigma}_q) + 0.5(\boldsymbol{\mu}_q - \boldsymbol{\mu}_p)^T \boldsymbol{\Sigma}_p^{-1}(\boldsymbol{\mu}_q - \boldsymbol{\mu}_p) - \frac{d}{2} \quad (E.11)$$

where  $|\Sigma_p|$  denotes the determinant of the matrix  $\Sigma_p$ .

## E.3 The Multinomial Distribution

If a random variable x can take one of one m discrete values  $x_1, x_2, .., x_m$  and

$$p(x = x_s) = \pi_s \tag{E.12}$$

then x is said to have a multinomial distribution.

# E.4 The Dirichlet Distribution

If  $\boldsymbol{\pi} = [\pi_1, \pi_2, ..., \pi_m]$  are the parameters of a multinomial distribution then

$$q(\boldsymbol{\pi}) = \Gamma(\lambda_{tot}) \prod_{s=1}^{m} \frac{\pi_s^{\lambda_s - 1}}{\Gamma(\lambda_s)}$$
(E.13)

defines a Dirichlet distribution over these parameters where

$$\lambda_{tot} = \sum_{s} \lambda_s \tag{E.14}$$

The mean value of  $\pi_s$  is  $\lambda_s/\lambda_{tot}$ .

### E.4.1 Relative Entropy

For Dirichlet densities  $q(\boldsymbol{\pi}) = D(\boldsymbol{\pi}; \boldsymbol{\lambda}_q)$  and  $p(\boldsymbol{\pi}) = D(\boldsymbol{\pi}; \boldsymbol{\lambda}_p)$  where the number of states is m and  $\boldsymbol{\lambda}_q = [\lambda_q(1), \lambda_q(2), ..., \lambda_q(m)]$  and  $\boldsymbol{\lambda}_p = [\lambda_p(1), \lambda_p(2), ..., \lambda_p(m)]$ . the KL-divergence is

$$D[q||p] = \Gamma(\log \lambda_{qtot}) + \sum_{s=1}^{m} (\lambda_q(s) - 1)(\Psi(\lambda_q(s)) - \Psi(\lambda_{qtot}) - \log \Gamma(\lambda_q(s)) E.15)$$
  
-  $\Gamma(\log \lambda_{ptot}) + \sum_{s=1}^{m} (\lambda_p(s) - 1)(\Psi(\lambda_q(s)) - \Psi(\lambda_{qtot}) - \log \Gamma(\lambda_p(s)))$ 

where

$$\lambda_{qtot} = \sum_{s=1}^{m} \lambda_q(s)$$

$$\lambda_{ptot} = \sum_{s=1}^{m} \lambda_p(s)$$
(E.16)

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and  $\Psi()$  is the digamma function.

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